

## 第5回クラスター階層領域研究会

# Ab initio description of light nuclei from no-core Monte Carlo shell model

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ZOOM

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# Nuclear Landscape

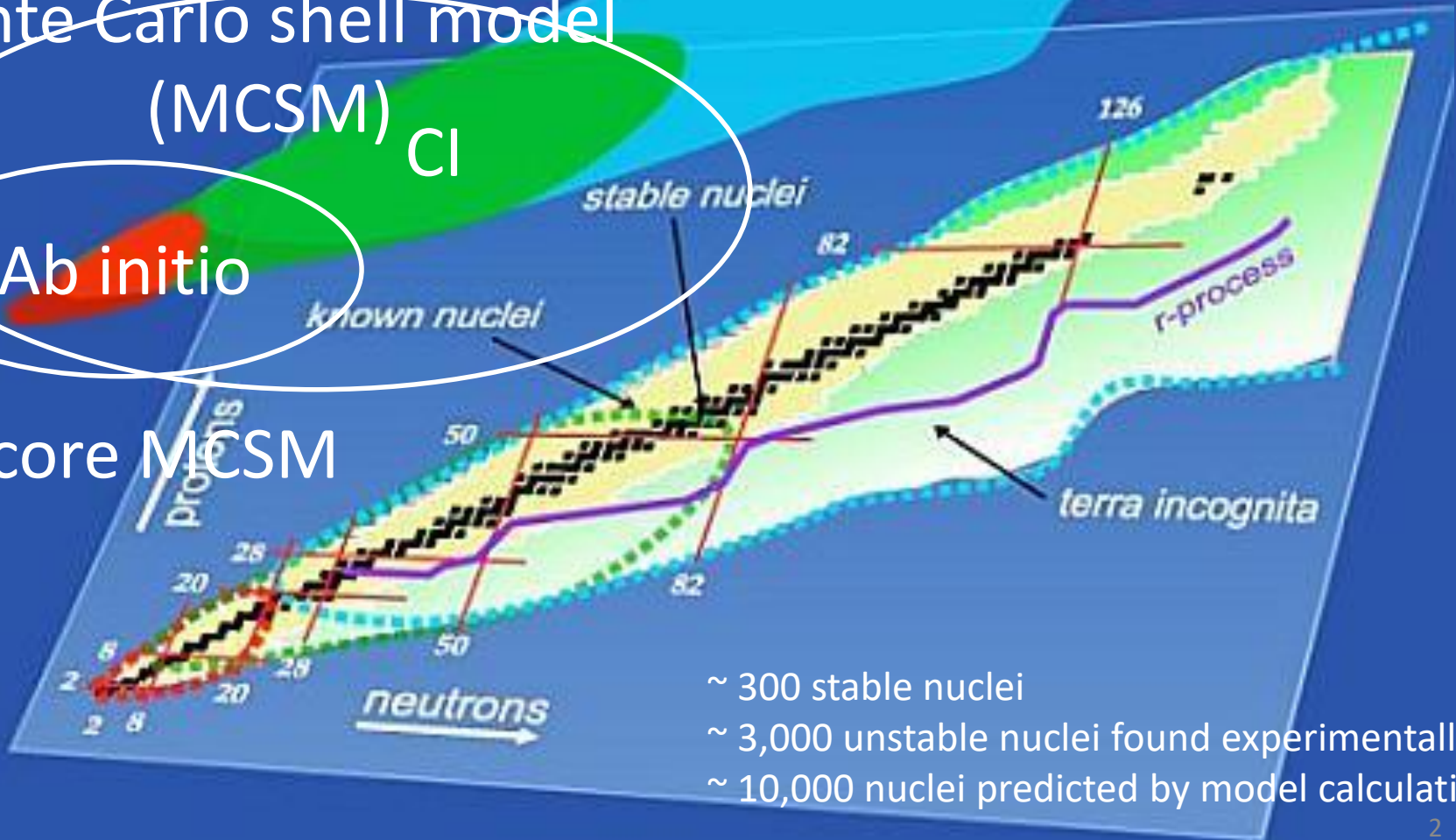


Monte Carlo shell model  
(MCSM) CI

Ab initio

No-core MCSM

DFT



- ~ 300 stable nuclei
- ~ 3,000 unstable nuclei found experimentally
- ~ 10,000 nuclei predicted by model calculations

# Configuration Interaction (CI)

- Eigenvalue problem of large and sparse Hamiltonian matrix

$$H|\Psi\rangle = E|\Psi\rangle$$

$$\begin{pmatrix} H_{11} & H_{12} & H_{13} & H_{14} & H_{15} & \cdots \\ H_{21} & H_{22} & H_{23} & H_{24} & & \\ H_{31} & H_{32} & H_{33} & & & \\ H_{41} & H_{33} & & \ddots & & \\ H_{51} & & & & & \\ \vdots & & & & & \end{pmatrix} \begin{pmatrix} \Psi_1 \\ \Psi_2 \\ \Psi_3 \\ \Psi_4 \\ \Psi_5 \\ \vdots \end{pmatrix} = \begin{pmatrix} E_1 & & & & & 0 \\ & E_2 & & & & \\ & & E_3 & & & \\ & & & \ddots & & \\ 0 & & & & & \end{pmatrix} \begin{pmatrix} \Psi_1 \\ \Psi_2 \\ \Psi_3 \\ \Psi_4 \\ \Psi_5 \\ \vdots \end{pmatrix}$$

Large and sparse real symmetric

$\sim \mathcal{O}(10^{11})$

# non-zero MEs  
 $\sim \mathcal{O}(10^{13-14})$

Lanczos method for diagonalization

Slater determinants

$$\left\{ \begin{array}{l} |\Psi_1\rangle = a_\alpha^\dagger a_\beta^\dagger a_\gamma^\dagger \cdots |-\rangle \\ |\Psi_2\rangle = a_\alpha^\dagger a_\beta^\dagger a_\gamma^\dagger \cdots |-\rangle \\ |\Psi_3\rangle = \cdots \\ \vdots \end{array} \right.$$

# Monte Carlo shell model (MCSM)

## Standard shell-model calculation

$$\mathbf{H} = \begin{pmatrix} * & * & * & * & * & \dots \\ * & * & * & * & & \\ * & * & * & & & \\ * & * & & \ddots & & \\ * & & & & \ddots & \\ \vdots & & & & & \ddots \end{pmatrix}$$

Diagonalization

$$\begin{pmatrix} E_0 & & & & & 0 \\ & E_1 & & & & \\ & & E_2 & & & \\ & & & \ddots & & \\ & & & & \ddots & \\ 0 & & & & & \end{pmatrix}$$

$$d \lesssim \mathcal{O}(10^{11})$$

Spanned by Slater determinants

## Monte Carlo shell model

Importance truncation

$$\mathbf{H} \sim \begin{pmatrix} * & * & \dots \\ * & \ddots & \\ \vdots & & \ddots \end{pmatrix}$$

Diagonalization

$$\begin{pmatrix} E'_0 & & 0 \\ & E'_1 & \\ 0 & & \ddots \end{pmatrix}$$

$$d_{\text{MCSM}} \sim \mathcal{O}(10^2)$$

Spanned by important bases  
selected stochastically and variationally

# MCSM wave function

- Superposition of quantum-number projected SDs

# of MCSM basis states ~ 100

$$|\Psi^{JM\pi}(N_b)\rangle = \underbrace{\sum_{d=1}^{N_b} f^{(d)}}_{\text{Superposition}} \underbrace{\sum_{K=-J}^J g_K^{(d)} \hat{P}^\pi \hat{P}_{MK}^J}_{\text{Projection on to good spin \& parity}} \underbrace{|\Phi(D^{(d)})\rangle}_{\text{Deformed SD}}$$

$$|\Phi(D^{(d)})\rangle = \prod_i \hat{a}_i^\dagger(D^{(d)}) | \text{core} \rangle \quad \hat{a}_i^\dagger(D^{(d)}) = \sum_\ell D_{li}^{(d)} \hat{c}_\ell^\dagger$$

|vacuum\rangle

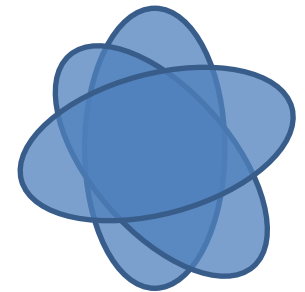
- Angular momentum projection

$$\hat{P}_{MK}^I = \frac{2I+1}{8\pi^2} \int P_{MK}^{I*}(\Omega) \hat{R}(\Omega) d\Omega$$

$\sum_{\Omega} W(\Omega) \dots$   
 $\Omega \quad O(\sim 10^4)$

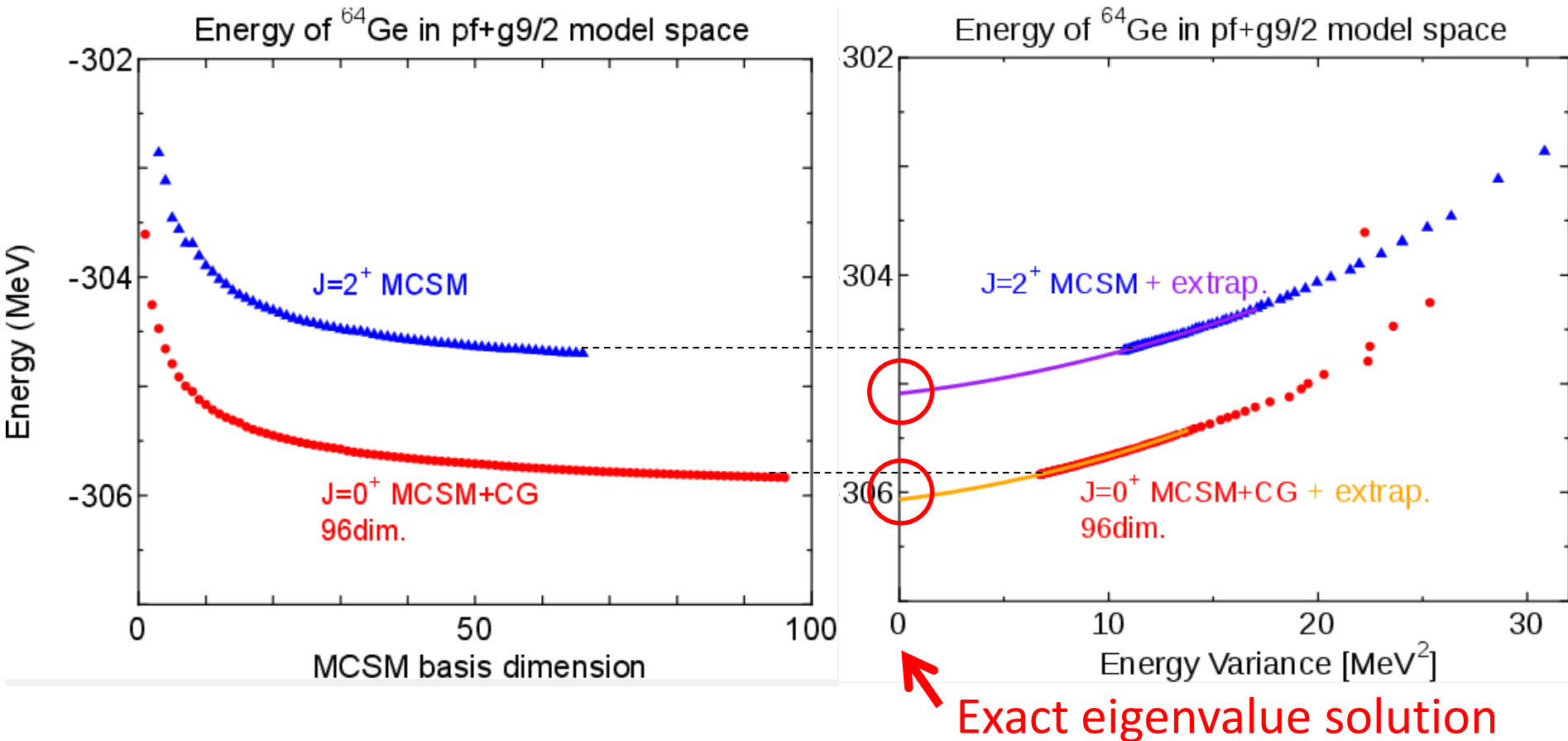
Restoration of symmetries

Favorable for  
massively parallel computation<sup>5</sup>

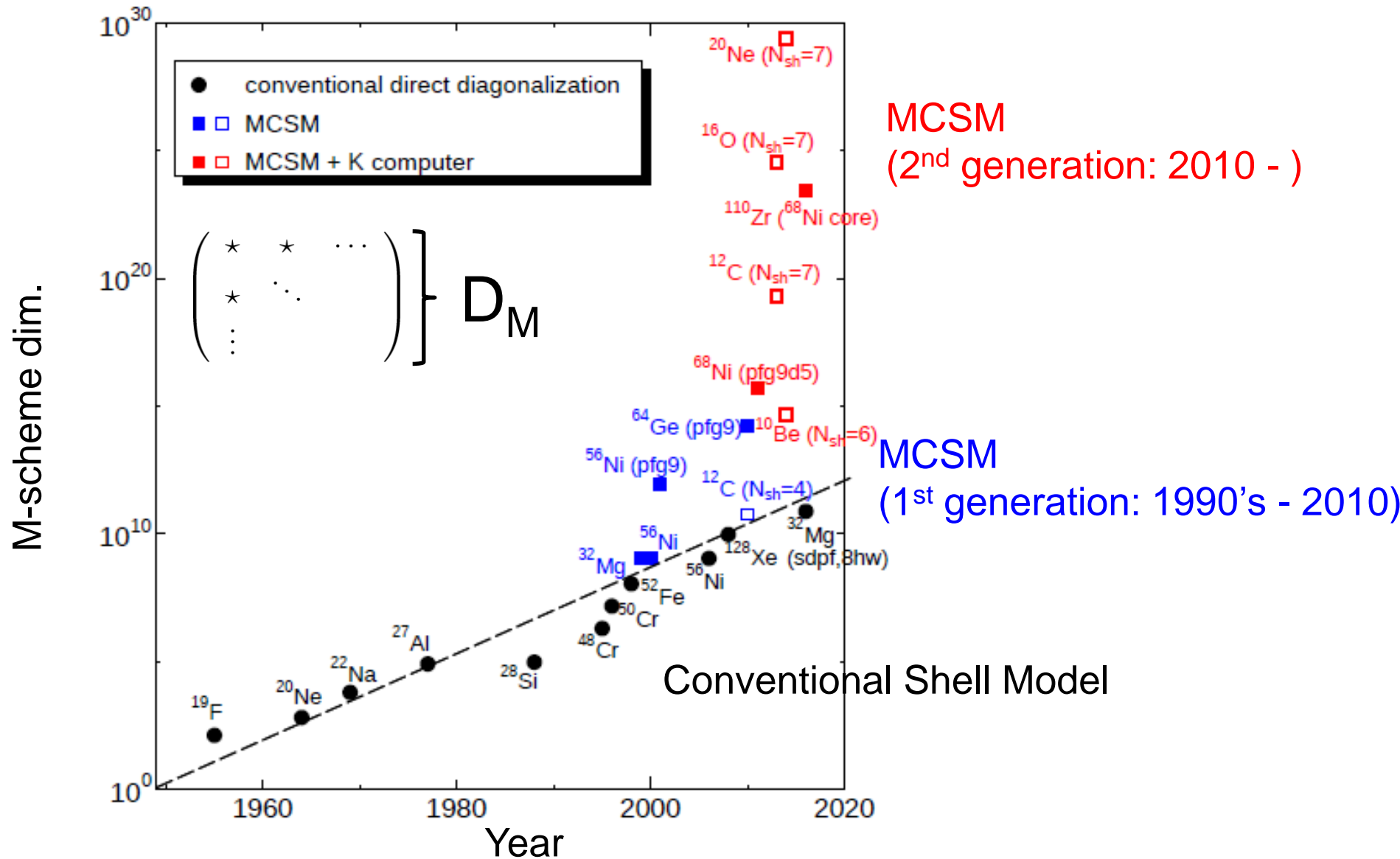


# Towards exact solution for eigenvalue

- Example for the energy-variance extrapolation in the MCSM  
 $^{64}\text{Ge}$  ( $Z=32, N=32$ ), pf+g9/2 model space ( $^{40}\text{Ca}$  core)



# Dimensionality of shell-model calculations



# “Ab initio” in low-energy nuclear structure physics

- Major challenge in nuclear physics
  - Nuclear structure & reactions directly from *ab-initio* calc. w/ nuclear forces
  - *ab-initio* approaches in nuclear structure calculations ( $A > 4$ ):
    - Light mass: Green’s Function Monte Carlo, No-Core Shell Model ( $A \sim 12$ ),
    - Medium/heavy mass: Coupled Cluster, IM-SRG,
    - Self-consistent Green’s Function theory, Lattice EFT, UMOA, ...
- Solve the non-relativistic many-body Schroedinger eq. and obtain the eigenvalues and eigenvectors.

$$H|\Psi\rangle = E|\Psi\rangle$$

$$H = T + V_{NN} + V_{3N} + \dots + V_{\text{Coulomb}}$$

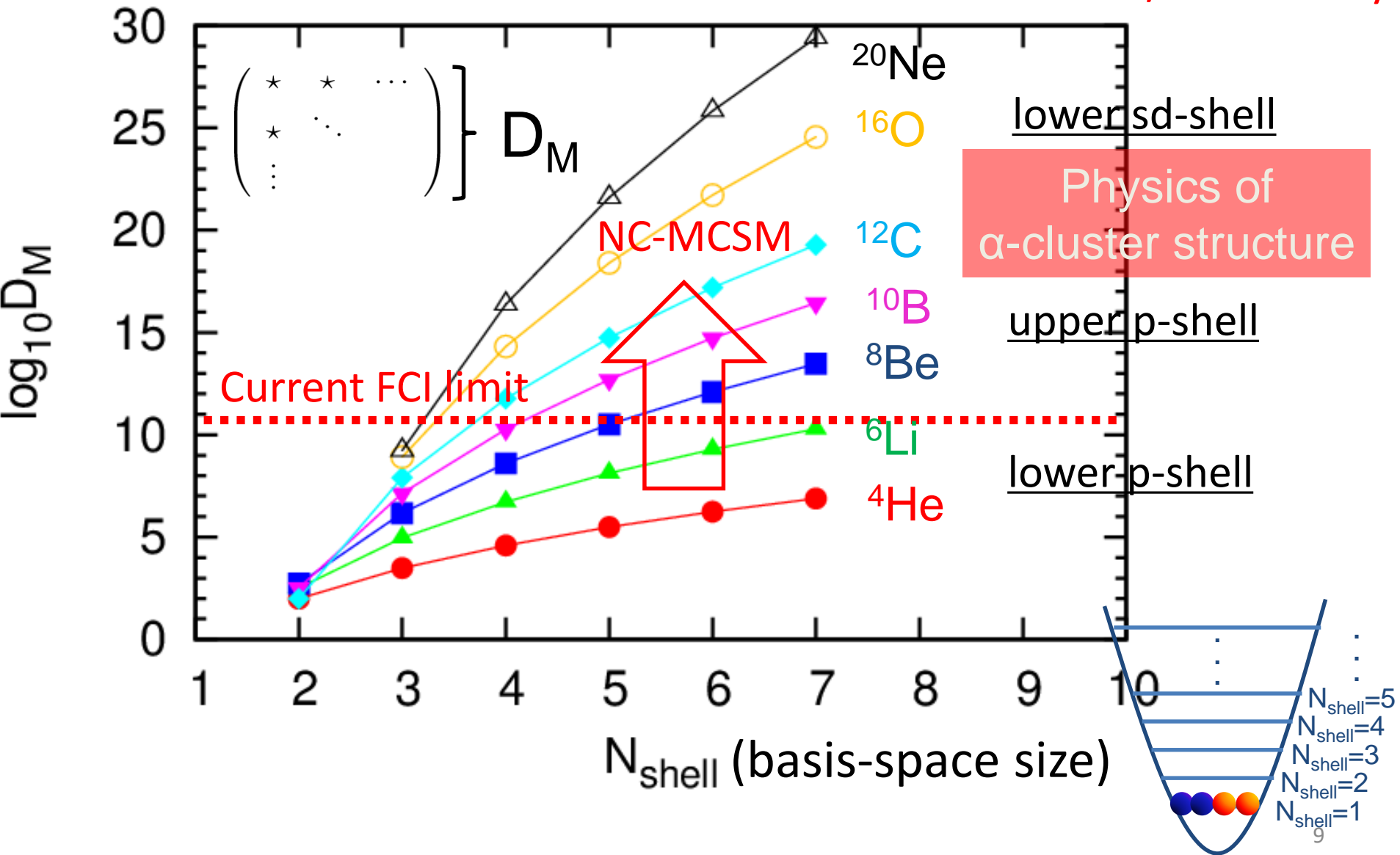
- *Ab initio*: All nucleons are active, and Hamiltonian consists of realistic NN (+ 3N + ...) potentials.

→ Computationally demanding → Monte Carlo shell model (MCSM)



# M-scheme dimension

Ab initio calculations w/ NN int only



# How to obtain ab-initio results from no-core MCSM

- Two steps of the extrapolation

← Same as in the MCSM w/ an inert core

1. Extrapolation of our MCSM (approx.) results to exact results in the finite size of model space

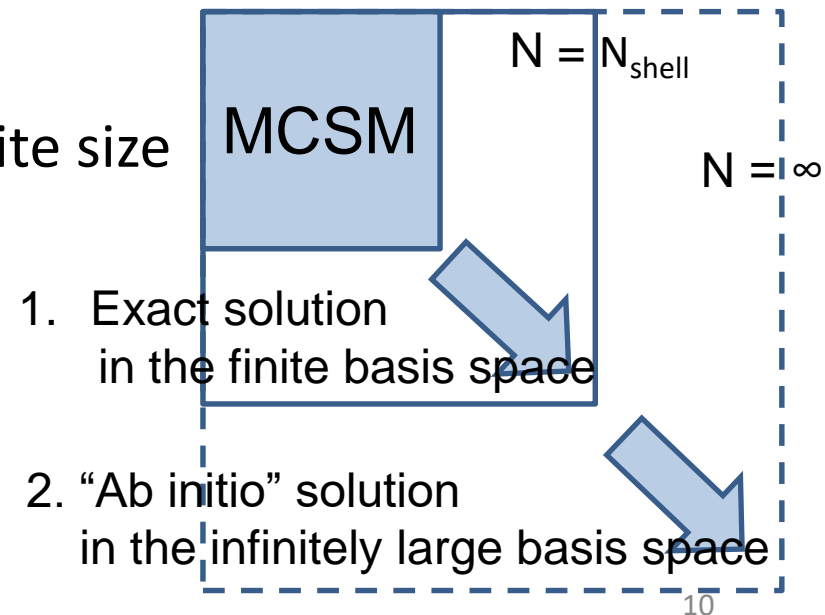
**Energy-variance extrapolation**

N. Shimizu, Y. Utsuno, T. Mizusaki, T. Otsuka, T. Abe, & M. Honma, Phys. Rev. C82, 061305(R) (2010)

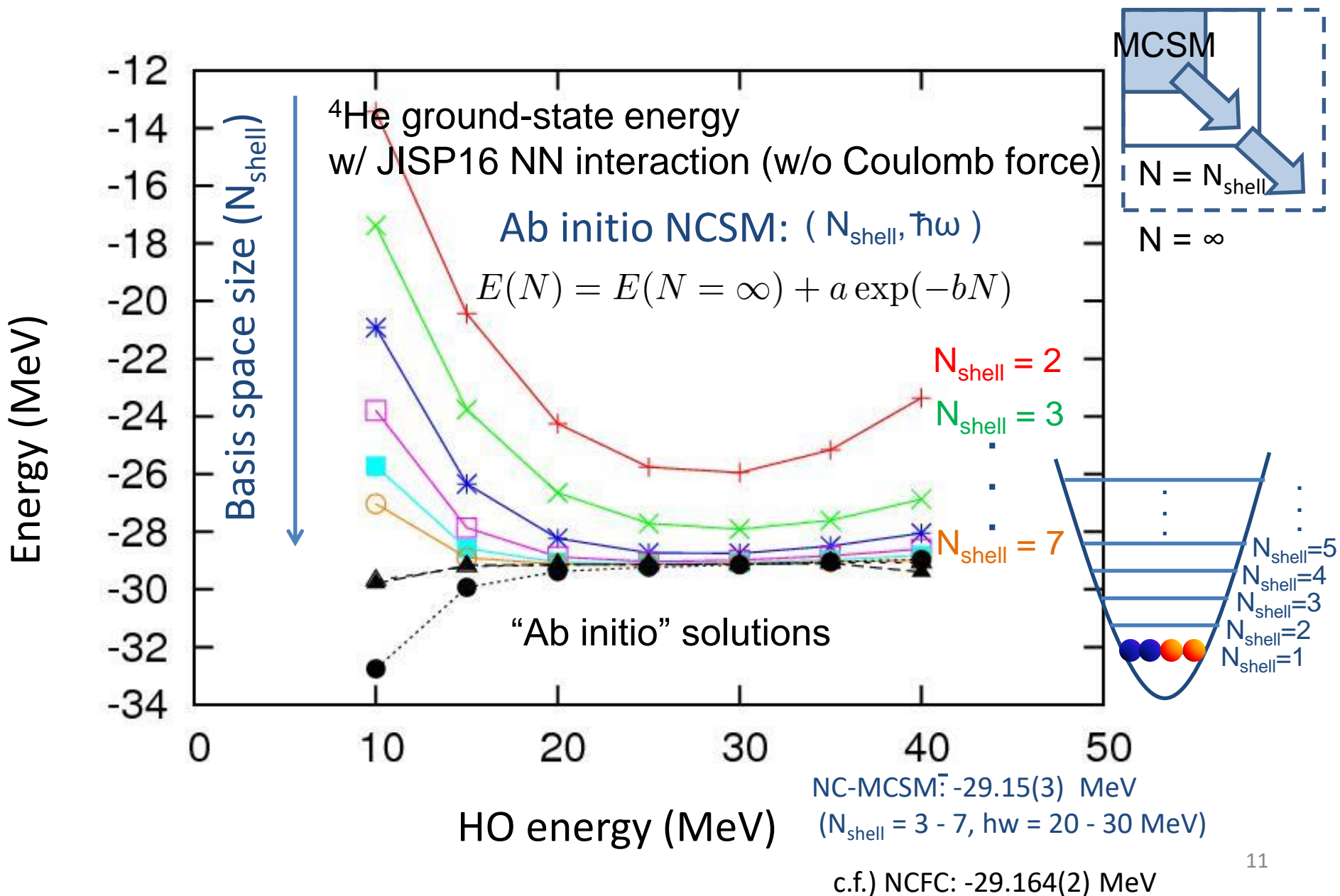
2. Extrapolation of the results in the finite size to the infinitely large basis space

- Empirical extrapolation w.r.t.  $N_{\text{shell}}$
- IR- & UV-cutoff extrapolations

→ **Ab initio solution**



# Extrapolation to the infinitely large basis space



# Inter-nucleon potentials

- JISP16:

J-matrix Inversion Scattering Potential tuned up to O-16

- Derived from nucleon-nucleon scattering phase shifts by J-matrix inversion scattering method.

Then, adjusted via a phase-shift equivalent transformations (PETs) to better describe light nuclei with  $A < 16$

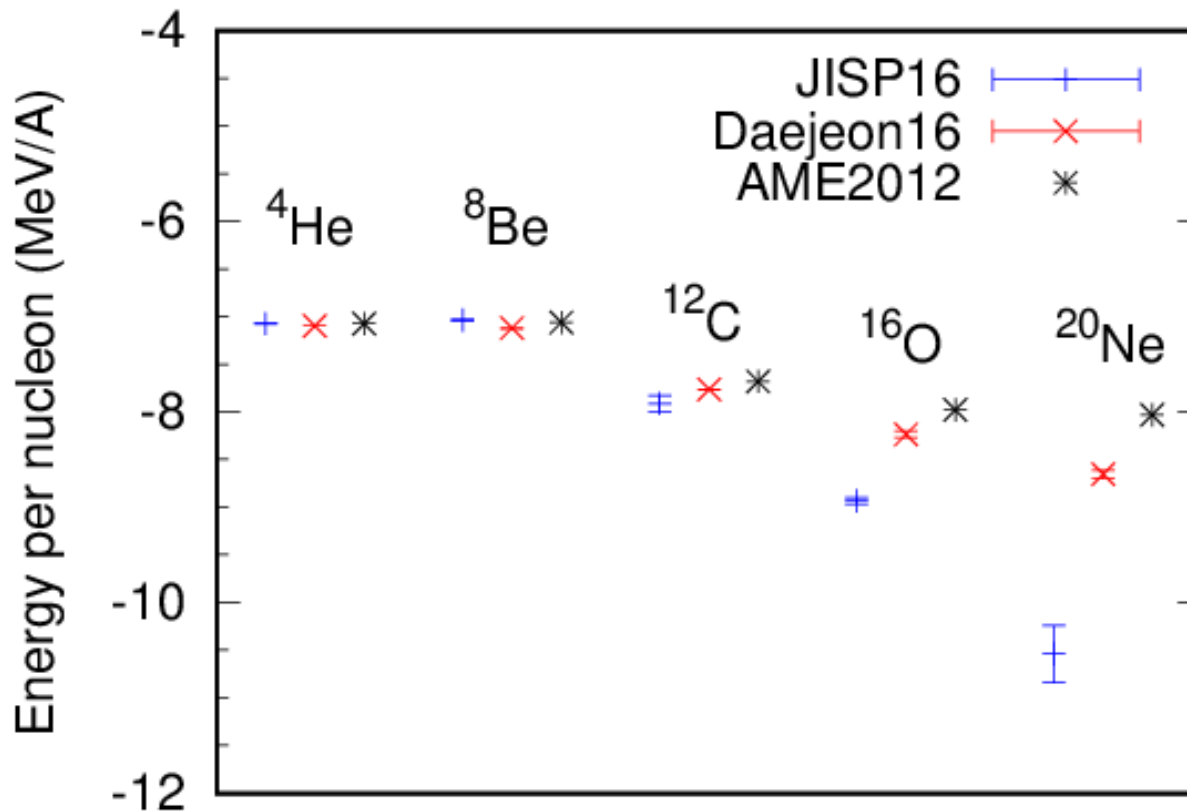
A. M. Shirokov, J. P. Vary, A. I. Mazur and T. A. Weber, PLB644, 33 (2007)

- Daejeon16:

- Starting from  $\chi$ EFT N3LO NN interaction (EM) + PETs

A. M. Shirokov, I. J. Shin, Y. Kim, M. Sosonkina, P. Maris and J. P. Vary, PLB761, 87 (2016).

# Ground-state energies of light nuclei

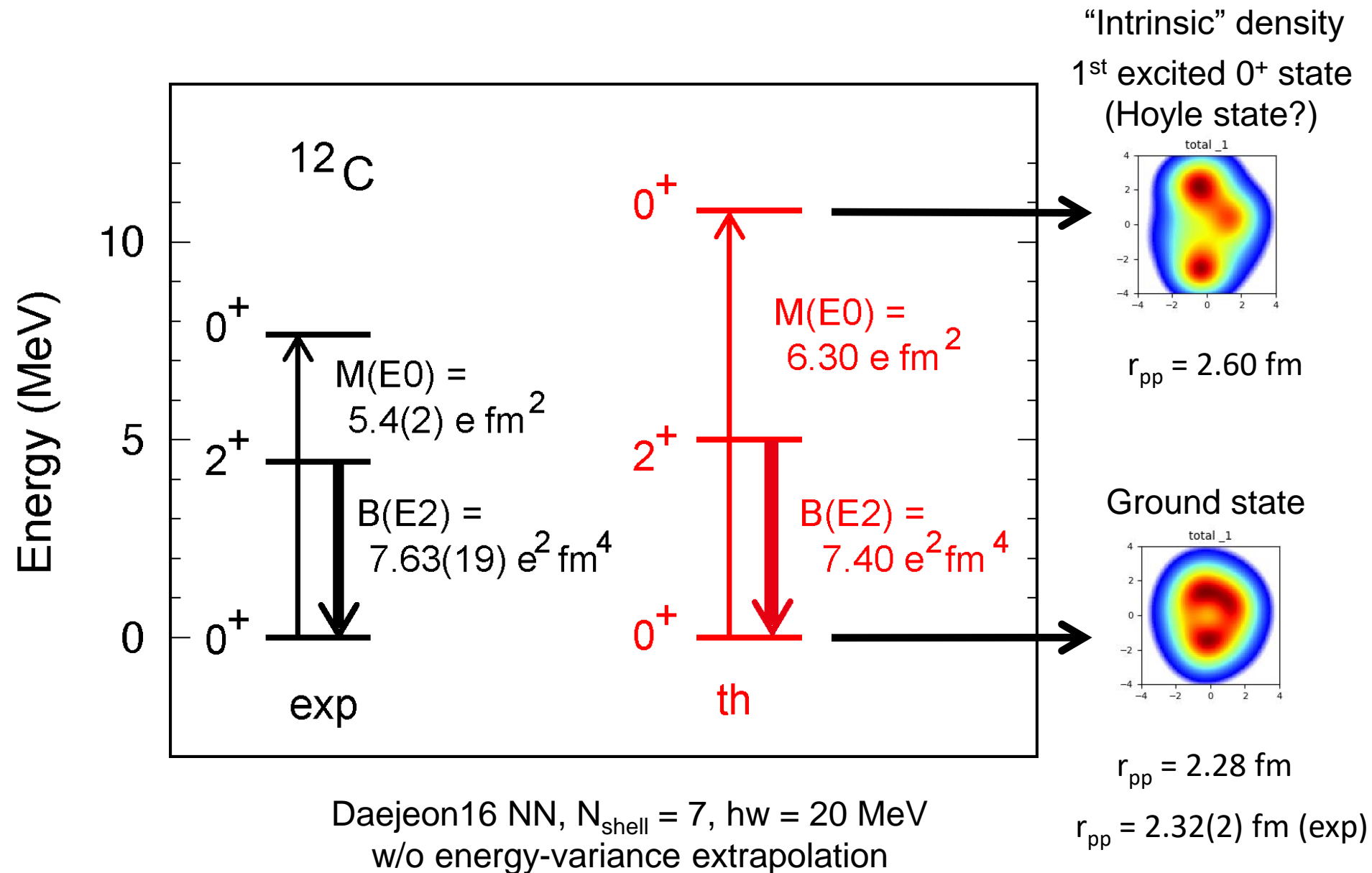


MCSM results are obtained using K computer by traditional extrapolation w/ optimum harmonic oscillator energies.

JISP16 results show good agreements w/ experimental data up to  $^{12}\text{C}$ , slightly overbound for  $^{16}\text{O}$ , and clearly overbound for  $^{20}\text{Ne}$ .

Daejeon16 results show good agreements w/ experimental data up to  $^{20}\text{Ne}$ .

# $^{12}\text{C}$ excitation spectra and transitions



# Summary

- Ab initio calculations by NC-MCSM can be performed in light nuclei ( $A \leq 20$ ).
- NC-MCSM results can be extrapolated to the infinitely large basis space to obtain ab initio solution.
  - Daejoen16 NN interaction provides good agreement w/ experimental data for light nuclei.

## Future perspective

- Alpha-cluster structure in light nuclei (p- & sd-shell region)
  - Next talk by T. Otsuka
- Neutron(proton)-rich nuclei in p-shell nuclei
- Heavier (sd-shell)  $N \sim Z$  nuclei beyond  $^{20}\text{Ne}$